

Chemical characteristics of chromian spinel in plutonic rocks: Implications for deep magma processes and discrimination of tectonic setting

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1 Chemical characteristics of chromian spinel in plutonic rocks:
2 implications for deep magma processes and discrimination of
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12 **Abstract** We summarize chemical characteristics of chromian spinels from ultramafic
13 to mafic plutonic rocks (lherzolites, harzburgites, dunites, wehrlites, troctolites and
14 olivine gabbros) with regard to three tectonic settings (mid-ocean ridge, arc and oceanic
15 hotspot). The chemical range of spinels is distinguishable between the three settings in
16 terms of Cr# (= Cr/(Cr + Al) atomic ratio) and Ti content. The relationships are almost
17 parallel with those of chromian spinels in volcanic rocks, but the Ti content is slightly

18 lower in plutonics than in volcanics at a given tectonic environment. The Cr# of spinels
19 in plutonic rocks is highly diverse; its ranges overlap between the three settings, but
20 extend to higher values (up to 0.8) in arc and oceanic hotspot environments. The Ti
21 content of spinels in plutonics increases, for a given lithology, from the arc to oceanic
22 hotspot settings via mid-ocean ridge on average. This chemical diversity is consistent
23 with that of erupted magmas from the three settings. If we systematically know the
24 chemistry of chromian spinels from a series of plutonic rocks, we can estimate their
25 tectonic environments of formation. The spinel chemistry is especially useful in dunitic
26 rocks, in which chromian spinel is the only discriminating mineral. Applying this,
27 discordant dunites cutting mantle peridotites were possible precipitated from arc-related
28 magmas in the Oman ophiolite, and from an intraplate tholeiite in the Lizard ophiolite,
29 Cornwall.

30

31 **Key words:** ultramafic plutonics, chromian spinel, tectonic setting, Ti content, Cr/(Cr +
32 Al) ratio

33 **Running title:** Chromian spinel in plutonic rocks

34

35

36 INTRODUCTION

37 Chromian spinel is common to ultramafic and related rocks, and is a very good indicator

38 of petrological characteristics of involved magmas (e.g. Irvine 1965, 1967; Dick &

39 Bullen 1984; Roeder 1994; Kamenetsky *et al.* 2001). It has a general formula, $(\text{Mg},$ 40 $\text{Fe}^{2+})(\text{Cr}, \text{Al}, \text{Fe}^{3+})_2\text{O}_4$, where Fe^{3+} is only minor in peridotitic rocks. $\text{Cr}/(\text{Cr} + \text{Al})$ atomic

41 ratio (= Cr#) is highly variable and serves as an important petrogenetic indicator for

42 ultramafic and related rocks (Irvine 1967; Dick & Bullen 1984). $\text{Mg}/(\text{Mg} + \text{Fe}^{2+})$ atomic

43 ratio (= Mg#) varies inversely with the Cr# in chromian spinel (e.g. Irvine 1967). Small

44 amounts of Ti are possibly incorporated as Fe_2TiO_4 (= ulvospinel component) in

45 chromian spinel. Arai (1992) summarized the chemistry of chromian spinel in volcanic

46 rocks (or magmas) as a potential indicator of magma chemistry for three main tectonic

47 settings, *i.e.*, the mid-ocean ridge, arc and intraplate. Irvine (1967) and Dick and Bullen

48 (1984) referred to the Cr# and Mg# of chromian spinel, and Arai (1994a, b) discussed

49 the relationship between the Cr# of chromian spinel and Fo of coexisting olivine in

50 peridotites and related rocks. The range of Cr# of chromian spinel alone, however, has

51 considerable overlaps between the different settings (Arai 1994a). The Mg# of chromian

52 spinel is strongly dependent on the equilibrium temperature (Irvine 1965; Jackson 1969),
53 and is changeable at subsolidus stage also depending on its modal amount (Arai 1980).
54 We should notice that the Mg#-Cr# relationship commonly used for descriptions and
55 discussions of chromian spinel in igneous rocks (Irvine 1967) is strongly dependent on
56 their subsolidus cooling histories after the igneous stage.

57 In this article, we review and summarize the chromian spinel chemistry (mainly the
58 Cr# and Ti content) in deep-seated rocks (lherzolites, harzburgites, dunites, wehrlites,
59 troctolites and gabbros) to define its chemical spread for discrimination of their tectonic
60 settings and deep magmatic processes. We examined chromian spinel compositions in
61 the ultramafic and mafic plutonic rocks (mantle peridotites and ultramafic/mafic
62 cumulates) for which derived tectonic settings are well constrained. Each of the three
63 settings, i.e., the mid-ocean ridge, arc (mantle wedge) and oceanic hotspot (intraplate),
64 produces the magmas that are distinguishable from those of the other settings in
65 geochemistry (e.g. Pearce 1975). There have been plenty of articles dealing with
66 ultramafic xenoliths from continental areas captured by intra-plate basalts. The
67 deep-seated rocks from non-arc continental areas are not used in this article, because
68 they may have complicated histories, i.e., multi-setting generation/modification, and are

69 not appropriate to the purpose of this study. The chromian spinel chemistry has been
70 systematically discussed both for mantle peridotites (e.g. Dick & Bullen 1984; Arai
71 1994a) and for volcanic rocks (e.g. Arai, 1992; Kamenetsky *et al.* 2001). No works has
72 ever discussed chromian spinel in possibly cumulative plutonic rocks (dunites, wehrlites,
73 troctolites and gabbros) in more systematic way than this article that deals with
74 chromian spinels from plutonic rocks only with well-constrained derivations. The result
75 of this work is potentially useful, because such rocks are quite commonly found from
76 various geologic bodies. We also present two examples of application of our result to
77 two ophiolitic dunites, which are apparently unknown or debated for the tectonic setting
78 of formation. This article is supplementary to Arai (1992), which deals with chromian
79 spinels in volcanic rocks from the three tectonic settings.

80

81 DATA ACQUISITION

82 We collected data of chromian spinel in deep-seated rocks from three main tectonic
83 settings, i.e., mid-ocean ridge, arc and oceanic hotspot. The source of spinel chemical
84 data treated here is unpublished theses of our laboratory in addition to the literature.
85 Fe^{2+} and Fe^{3+} amounts were calculated assuming spinel stoichiometry. Titanium is

86 assumed to form the ulvospinel component.

87

88 MID-OCEAN RIDGES

89 We can obtain deep-seated rocks (lherzolites, harzburgites, dunites, wehrlites, troctolites

90 and gabbros) from the present-day ocean floor by dredging, drilling and submersible

91 diving (e.g. Dick 1989). The oceanic fracture zones (FZ), which are more prominently

92 developed in the slow to ultraslow spreading ridge systems than in fast spreading ones,

93 are the main loci for obtaining abyssal deep-seated rocks. Hess Deep, the East Pacific

94 Rise, is one of the non-FZ localities where deep-seated rocks are exposed on the ocean

95 floor (e.g. Arai & Matsukage 1996; Dick & Natland 1996; Allan & Dick 1996). We can

96 interpret these plutonics as deep-seated magmatic products beneath the mid-ocean ridge.

97 Harzburgites and lherzolites are predominant in the uppermost mantle of fast spreading

98 ridges and of slow spreading ridges, respectively (Niu & Hékinian 1997). Dunites and

99 related rocks (troctolites and olivine gabbros) are commonly found from Hess Deep,

100 and they may represent the Moho transition zone of the fast spreading ridge (e.g. Arai &

101 Matsukage 1996).

102

103 OPHIOLITES

104 Plutonic rocks from ophiolites may represent the deep-seated rocks of some sorts of
105 oceanic lithosphere (e.g. Coleman 1977; Nicolas 1989). Their tectonic setting for
106 genesis has been a problem in controversy since the pioneering paper of Miyashiro
107 (1973), and therefore, the data from ophiolitic plutonic rocks are not considered in this
108 section. Some ophiolites that exhibit both mid-ocean ridge and island-arc characteristics
109 are called “supra-subduction zone” (SSZ) ophiolites (Pearce *et al.* 1984). Many people
110 have favored the back-arc basin as the locus of the SSZ ophiolite formation (e.g. Pearce
111 *et al.* 1984; Moores *et al.* 1984). The polygenetic nature of some ophiolites has been
112 recently recognized as well. For example, some peridotites from the northern Oman
113 ophiolite are of arc origin (e.g. Tamura & Arai 2006; Arai *et al.* 2006), although the
114 main portion of the mantle section was of mid-ocean ridge origin (e.g. Nicolas 1989).
115 Wehrlitic rocks around the Moho transition zone were interpreted as mid-ocean ridge
116 products in the southern Oman ophiolite (Koga *et al.* 2001). For another example, the
117 mantle member of the Coast Range ophiolite, California, is composed of a mixture of
118 SSZ harzburgites and abyssal lherzolites (Choi *et al.* 2008; Jean *et al.* 2010).

119

120 ARCS

121 Deep-seated rocks from the sub-arc mantle (mantle wedge) are more difficult to obtain
122 systematically. Alkali basalts that carry deep-seated rocks as xenoliths most frequently
123 erupt on non-arc regions, i.e., on continental rift zones or oceanic hotspots, and the
124 xenoliths in kimberlites represent the upper mantle beneath cratons (e.g. Nixon 1987).
125 Genesis of calc-alkaline magmas is related with the subduction of slab (e.g. Tatsumi &
126 Eggins 1995), and their deep-seated xenoliths undoubtedly represent the sub-arc
127 deep-seated materials. Calc-alkaline andesites and basaltic andesites from Megata and
128 Oshima-Ôshima volcanoes (Northeast Japan arc), Iraya volcano (Luzon arc), and
129 Avacha and Shiveluch volcanoes (Kamchatka arc) contain peridotite xenoliths that are
130 derived from lithosphere of the mantle wedge (e.g. Takahashi 1978; Ninomiya & Arai
131 1992; Arai *et al.* 2003, 2004; Ishimaru *et al.* 2007; Bryant *et al.* 2007). It is noteworthy
132 that some of ultramafic rocks treated here form composite xenoliths with gabbros and
133 hornblendites in calc-alkaline volcanics (e.g. Ninomiya & Arai 1992; Arai *et al.* 1996,
134 2003, 2004). A variety of peridotites, i.e. lherzolites to highly refractory harzburgites
135 (Cr# of spinel < 0.8), may constitute the mantle wedge (Arai 1994a; Arai *et al.* 2003;
136 Arai & Ishimaru 2008). Dunites and related rocks form a thick cumulus mantle beneath

137 the Southwest Japan arc (Takahashi 1978). Dunites are relatively abundant from the
138 Oshima-Ôshima volcano (e.g. Yamamoto 1984), and are small in amount in other
139 localities (especially the Megata, Iraya and Avacha volcanoes).

140 Although not treated here, we can obtain large amounts of ultramafic xenoliths
141 carried by non-arc type alkaline basalts on past or present-day arcs, e.g., the Japan arcs
142 (e.g. Takahashi 1978; Aoki 1987; Abe *et al.* 1998, 1999; Arai *et al.* 1998, 2000, 2007).

143 These materials from the Japan arcs may represent sub-arc mantle materials because
144 their eruption ages are mostly younger than Miocene (Uto 1990), when the present arc
145 setting had been established for the Japan island arcs (e.g. Otofuji *et al.* 1985).

146 Ultramafic to mafic xenoliths from the Southwest Japan arc have been affected to
147 various extents (Arai *et al.* 2000) by Cenozoic non-arc type alkali basalt magmas
148 (Nakamura *et al.* 1987).

149 Peridotites and related rocks are exposed on the ocean floor of fore-arc regions (e.g.
150 Fisher & Engel 1969), and the dredged and drilled peridotites distinctively represent
151 fore-arc mantle materials (e.g. Bloomer 1983; Parkinson & Pearce 1998). Refractory
152 harzburgites are dominant in the fore-arc mantle (Arai 1994a). In this article, we treat
153 only these two sets of deep-seated rocks, i.e., the ultramafic xenoliths in arc-type

154 volcanics and the ultramafic rocks exposed on the present-day fore-arc ocean floor, as
155 “genuine” sub-arc materials with well-defined derivations.

156

157 OCEANIC HOTSPOTS

158 Deep-seated rocks from the oceanic hotspot areas have been almost solely obtained as
159 ultramafic and mafic xenoliths in their volcanic rocks (e.g. Nixon 1987). The xenoliths
160 in alkaline basalts especially represent the lower crust to upper mantle of that tectonic
161 setting because the xenolith-bearing magmas postdate the main stage of hotspot
162 volcanism (e.g. Jackson & Wright 1970). Lherzolites are apparently dominant in amount
163 as the upper mantle material from the oceanic hotspot (Arai 1994a). Some peridotites
164 may be related with the hotspot magmatism (cumulates or restites), and the others are
165 only representative of the sub-oceanic mantle that hosts mantle plumes relevant to the
166 hotspot activity. Arai (1994b) predicted predominance of refractory harzburgites with
167 high Cr# (around 0.7) of chromian spinel as residual peridotites after the hotspot
168 tholeiite genesis. Some basalt magmas contain a large amount of dunite xenoliths,
169 indicating a thick dunite layer at the uppermost mantle as a result of extensive tholeiitic
170 magmatism (Jackson & Wright 1970; Sen & Presnall 1986). Hawaiian and French

171 Polynesian hotspots are especially important for occurrences of ultramafic xenoliths

172 (Nixon 1987), and our data accumulation owes the literature dealing with them.

173

174 CHEMICAL SPECTRA OF CHROMIAN SPINELS IN PLUTONIC

175 ROCKS FROM THE THREE TECTONIC SETTINGS

176

177 MID-OCEAN RIDGES

178 As is well known, the Cr# of chromian spinel in abyssal mantle peridotites ranges from

179 0.1 to 0.6 (e.g. Dick & Bullen 1984; Arai 1994a; Niu & Hékinian 1997) (Fig. 1a). It

180 changes from around 0.4 to 0.6 in harzburgites to <0.4 in lherzolites, in response to a

181 decrease of degrees of partial melting (e.g. Dick & Bullen, 1984; Arai 1994a). The TiO₂

182 content is mostly lower than 0.3 wt% in their chromian spinel (Fig. 1a). The chromian

183 spinel interestingly displays the same range of Cr# between plagioclase-bearing and

184 –free varieties of mantle peridotites (Fig. 1a). The TiO₂ content of chromian spinel is,

185 however, systematically higher in the plagioclase-bearing peridotites than in the

186 plagioclase-free ones (Dick 1989). The chromian spinel exhibits relatively wide ranges

187 of Cr#, 0.2-0.6, and TiO₂, nil to 2 wt% (mostly <1 wt%) in abyssal dunites (Fig. 1a).

188 Other plutonic rocks, especially troctolites and olivine gabbros from Hess Deep, show a
189 narrow range of Cr#, mostly 0.5 to 0.6, and a wide range of TiO₂ content, <3 wt% in
190 chromian spinel. Chromian spinel in abyssal plutonic rocks is characterized by overall
191 low Fe³⁺ contents as that in MORB (Arai 1992). The Mg# is negatively correlated with
192 the Cr# in chromian spinel for all peridotitic rocks including dunites from the mid-ocean
193 ridges (Fig. 3a). Y_{Fe} (= Fe³⁺/(Cr + Al + Fe³⁺) atomic ratio) of chromian spinel is mostly
194 lower than 0.1 in peridotitic rocks, and lower than 0.2 in troctolites and gabbros (Figs.
195 2a and 4a). Chromian spinel shows lower Mg#s at a Cr# around 0.5 in abyssal gabbros
196 and troctolites. The TiO₂ content is well correlated positively with the Y_{Fe} in chromian
197 spinel from troctolites and gabbros, being 2.5 to 3 wt% at around Y_{Fe} of 0.2 (Fig. 4a).

198 Rocks of dunite-troctolite-olivine gabbro suite from Hess Deep, East Pacific Rise,
199 were interpreted as a reaction product between the primary MORB and mantle
200 harzburgite (Arai & Matsukage 1996; Dick & Natland 1996). These rocks are expected
201 to be in equilibrium with MORB in terms of mineral chemistry (e.g. Kelemen *et al.*
202 1995; Arai 2005). Plagioclase in the harzburgites/lherzolites is calcic, and is a
203 melt-impregnation product (e.g. Dick 1989). The formation of plagioclase-bearing
204 peridotites is the very initiation of melt/peridotite reaction. The slightly but clearly

205 higher TiO₂ content of chromian spinel in plagioclase-bearing peridotites (Fig. 1a) is
206 consistent with this interpretation (e.g. Dick 1989).

207

208 ARCS

209 The Cr# of chromian spinel exhibits a wide range, from than <0.2 to 0.9, for mantle
210 peridotites (lherzolite to harzburgite) and dunites (Fig. 1b). This is consistent with the
211 wide range of spinel Cr# in sub-arc mantle restites estimated from arc magmas (Arai
212 1994b). The TiO₂ content of chromian spinel is, however, slightly higher in dunites than
213 in mantle peridotites (Fig. 1b). Some of dunite, wehrlite and clinopyroxenite treated
214 here possibly have initially formed composite xenoliths with younger gabbros or
215 hornblendites, and have been chemically influenced by evolved magmas that formed the
216 latter younger rocks (e.g. Ninomiya & Arai 1992). The relatively high contents of TiO₂
217 and Fe³⁺ of some plutonic spinels are possibly due to such a secondary effect. Almost all
218 sub-arc spinels have low values of Y_{Fe}, < 0.3 (Fig. 2b). As is well known, the Mg#
219 shows roughly negative correlations with the Cr# (e.g. Irvine 1967; Dick & Bullen
220 1984) (Fig. 3b). Two Mg#-Cr# spinel trends can be recognized in mantle peridotites,
221 especially harzburgites, corresponding to two different derivations of the samples

222 treated here, namely the fore-arc rocks and xenoliths in arc magmas (Fig. 3b). This is
223 due to the difference of equilibrium temperature between the two rock suites (e.g.
224 Okamura *et al.* 2006), resulting from a decrease of Mg# of chromian spinel with
225 decreasing the equilibrium temperature in peridotites (Irvine 1967; Evans & Frost 1975).
226 Chromian spinel in some dunites, wehrlites and clinopyroxenites shows lower Mg#s at
227 given Cr#s (Fig. 3b). The TiO₂ content is positively correlated with the Fe³⁺ ratio for the
228 main cluster of sub-arc spinels, being 1 to 2 wt% at Y_{Fe} = 0.2 (Fig. 4b).

229

230 OCEANIC HOTSPOTS (PLUMES)

231 Ultramafic xenoliths have been extensively described from various oceanic hotspots on
232 the Earth, especially from Hawaii and the French Polynesian (e.g. Nixon 1987). The
233 Cr# of chromian spinel also changes from 0.1 to 0.8 with a lithological change from
234 lherzolite to harzburgite (Fig. 1c). The TiO₂ content of the peridotite spinel is mostly
235 lower than 4 wt% (Fig. 1c). The Cr# of chromian spinel shows almost the same range
236 between the mantle peridotites and dunites. The TiO₂ content of spinel is generally
237 higher in dunites than in mantle peridotites, and show the highest values, up to > 10
238 wt%, at the Cr# around 0.5 to 0.6 (Fig. 1c). Most of wehrlite spinel have relatively high

239 Cr#, around 0.6, and TiO₂ content, up to 6 wt% (Fig. 1c). The Y_{Fe} of chromian spinel is
240 highest around the intermediate Cr#, 0.5 to 0.6, and positively correlated to the TiO₂
241 content (Figs. 2c and 4)c. The TiO₂ content of hotspot spinels varies at a given Y_{Fe},
242 ranging from 1 to 6 wt% at Y_{Fe} = 0.2 (Fig. 4c). Harzburgite spinels have higher Mg# at
243 a given Cr# than dunite ones (Fig. 3c). As in the case of abyssal plutonic rocks, the Mg#
244 is extended toward lower values at the highest Cr# of the whole range, 0.6 to 0.7, in the
245 dunite spinels (Fig. 3c).

246

247 DISCUSSION

248

249 DISTINCTION OF THE THREE TECTONIC SETTINGS

250 Apart from the mantle peridotite, the plutonic rocks that bear chromian spinel are
251 mainly dunite, troctolite and olivine gabbro from the ocean floor, but are dunite and
252 wehrlite from the arc and the hotspot (Figs. 1 to 4). This indicates that the phase
253 crystallizing next to olivine is mainly plagioclase in MORB but clinopyroxene in both
254 arc and intraplate magmas. This is in turn related with the degree of partial melting in
255 the mantle, which is lower, on average, in the mid-ocean ridge than in sub-arc and in

256 hotspot conditions (e.g. Arai 1994a,b).

257 The Cr# ranges of chromian spinel in plutonic rocks are overlapping with each
258 other around 0.1 to 0.6 for the three tectonic settings, i.e., the mid-ocean ridge, arc and
259 intraplate (Fig. 1). It is difficult, therefore, to distinguish the tectonic settings in terms of
260 Cr# of spinel alone. The Cr# of spinel is barely higher than 0.6 in plutonic rocks from
261 the mid-oceanic ridges. It is frequently over 0.6, and is up to 0.9 for the arc setting, and
262 up to over 0.7 for the intraplate (hotspot or plume) setting. The TiO₂ content in
263 chromian spinel combined with the Cr# is, however, convenient for distinction of
264 plutonic rocks between the three tectonic settings (Fig. 1). The TiO₂ content of
265 chromian spinel in plutonic rocks decreases on average from the intraplate setting to arc
266 via mid-ocean ridge setting (Fig. 1). It is concluded that deep-seated ultramafic rocks
267 can be distinguished as a group with each other in terms of spinel chemistry, especially
268 Cr# and Ti content (Fig. 1). This distinction is consistent with the diversity of chromian
269 spinel in volcanics depending on the tectonic setting (Arai 1992).

270

271 IMPLICATIONS FOR DEEP MAGMATIC PROCESSES

272 All kinds of plutonic rock have relatively low-Ti spinels from the arc setting, and

273 dunites are almost indistinguishable from harzburgites (or lherzolites) in terms of spinel
274 chemistry (see Arai 1994b). Chromian spinel in dunites, troctolites and
275 melt-impregnated harzburgites (plagioclase harzburgites) from the ocean floor is high
276 both in Cr# (around 0.6 to 0.7) and in TiO₂ (Fig. 1a). Dunites and wehrlites from the
277 oceanic hotspot also contain high-Cr# and high-Ti chromian spinels (Fig. 1c). The wide
278 range of TiO₂ content at a given Y_{Fe} for hotspot dunite spinels (Figs. 1 and 4) is possibly
279 due to a variety of dunites, from those related with older MORB genesis to those related
280 to younger hotspot magmatism. It is noteworthy that the hotspot plutonic spinels are
281 lower in Ti at a given Y_{Fe} than the mid-ocean ridge ones (Fig. 4), despite that the
282 relations are the reverse for volcanic spinels (Arai 1992). Abyssal plutonic rocks are
283 mostly troctolites (Fig. 4a), in which Ti and Fe³⁺ are partitioned to chromian spinel
284 because plagioclase is free of these components. In contrast, Ti and Fe³⁺ are partitioned
285 to both clinopyroxene and chromian spinel in wehrlites, which are common from
286 hotspot environments (Fig. 4c). This is also related with the redox condition;
287 deep-seated magmas are more oxidized for the hotspot environments than for the
288 mid-ocean ridge ones. This is concordant to the difference of oxidation states between
289 the hotspot magmas and MORB (e.g. Christie *et al.* 1986; Rhodes & Vollinger 2005).

290 Discrepancy in spinel chemistry between effusive rocks and related plutonic rocks
291 is sometimes noticeable; volcanic spinels are sometimes more limited in Ti and Y_{Fe}
292 ranges than plutonic spinels (Fig. 4). This is primarily due to effective magmatic
293 evolution to concentrate these components within closed melt pools in deep parts (Arai
294 *et al.* 1997). Difference in spinel chemistry is striking between MORB and abyssal
295 plutonics (dunites, troctolites to olivine gabbros) (Fig. 4a). The TiO_2 and Y_{Fe} of
296 chromian spinel are limited, mostly <1 wt% and <0.1, respectively in MORB (Arai
297 1992), as compared to the values in abyssal plutonics (Fig. 4a).

298 It is noteworthy that high-Ti chromian spinels are also high in Cr# (around 0.6 to
299 0.7) (Fig. 1). The high-Cr#, -Ti spinels are common to dunites and related rocks from
300 the ocean floor and oceanic hotspot, part of which have been thought to be
301 peridotite/magma reaction products (e.g. Arai & Matsukage 1996; Dick & Natland
302 1996). Some of dunite and wehrlite xenoliths from the island arc setting are also
303 reaction products (e.g. Arai & Abe 1994), but the concerned arc magmas, which are
304 initially low-Ti, have not increased the Ti contents of chromian spinel even through the
305 peridotite/melt reaction processes.

306

307 EXAMPLES OF APPLICATION TO OPHIOLITES

308 The origin and nature of ophiolites have been controversial (e.g. Pearce *et al.* 1984;
309 Nicolas 1989). Arai *et al.* (2006), for example, suggested polygenetic nature of the
310 mantle part of the northern Oman ophiolite. We show two examples of discordant
311 dunites from two ophiolites as below. The dunite is an important constituent of the
312 Moho transition zone to upper mantle section of ophiolites, but the mineralogy is too
313 simple to constrain its derivation. If we apply the systematics discussed here, the
314 chromian spinel is indicative of the tectonic setting of the dunite formation.

315

316 DISCORDANT DUNITES FROM THE MANTLE SECTION OF THE

317 NORTHERN OMAN OPHIOLITE

318 It has been well recognized that the mantle section of the Oman ophiolite is dominated
319 by harzburgites (e.g. Boudier & Coleman 1981; Lippard *et al.* 1986). Lherzolites are
320 absent except at the base of the ophiolite (e.g. Lippard *et al.* 1986; Takazawa *et al.*
321 2003). The harzburgites mainly constitute the mantle section, containing spinels with
322 Cr#s <0.6 (Le Mée *et al.* 2004), similar to those obtained from the ocean floor of fast
323 spreading ridge origin (Niu & Hekinian 1997). Tamura and Arai (2006) found a

324 harzburgite-orthopyroxenite-dunite suite of sub-arc chemical affinity from the northern
325 Oman ophiolite. Arai *et al.* (2006) examined chemical variations of detrital chromian
326 spinel particles derived from the mantle section from recent riverbeds in the Oman
327 ophiolite. They found more than 20 to 30 percent of the total detrital chromian spinel
328 grains examined have Cr#s higher than 0.6 (Arai *et al.* 2006).

329 Discordant dunites cutting foliated harzburgites are very common in the mantle
330 section (Fig. 5a). They form dikes or networks, and some of them contain chromian
331 spinel concentrations (podiform chromitites) (e.g. Augé 1987; Ahmed & Arai 2002).
332 They are massive in appearance and solely comprise olivine and euhedral to subhedral
333 chromian spinel (Fig. 5b). We examined chromian spinels in the discordant dunites from
334 Wadi Rajmi and Wadi Fizh areas of the northern Oman ophiolite (Fig. 5a). The Cr# of
335 chromian spinel ranges from 0.4 to 0.8 with very low amounts of TiO₂, mostly < 0.3
336 wt% (Fig. 6). Olivine associated with the chromian spinel is around Fo₉₀₋₉₂ in
337 composition. The Oman discordant dunites are most likely to have been related with arc
338 magmas (see Figs. 1, 2 and 4). The mantle section of the Oman ophiolite, therefore,
339 comprises the ocean-floor peridotites (mainly harzburgites) modified by addition of
340 dunites of sub-arc affinity. This suggests a switch of tectonic setting from mid-ocean

341 ridge to arc (= supra-subduction zone) for genesis of the Oman ophiolite (Arai *et al.*
342 2006). Alternatively, this characteristic can be obtained at a back-arc tectonic setting,
343 where various magmas, from MORB-like to arc-type, are available (e.g. Pearce *et al.*
344 1984).

345

346 DISCORDANT DUNITES FROM THE LIZARD OPHIOLITE, CORNWALL

347 The mantle section of the Lizard ophiolite, Cornwall (Kirby 1979), is mainly composed
348 of lherzolites and concordant dunites (Green 1964; Kadoshima & Arai 2001). This is
349 very similar to a peridotite suite from the ocean floor of slow spreading ridge origin
350 (Roberts *et al.* 1993) if we consider the abundance of lherzolite over harzburgite (e.g.
351 Niu & Hekinian 1997). The Cr# ranges from 0.1 to 0.5 for the Lizard detrital spinels
352 (Kadoshima & Arai 2001), exactly being the same as those of abyssal peridotites
353 (lherzolites to harzburgites) (Fig. 6). This is consistent with the idea that the Lizard
354 peridotite was representative of the uppermost sub-oceanic mantle of a slow spreading
355 ridge, which may be composed of predominant lherzolites and subordinate harzburgites
356 (Arai 2005).

357 Networks of younger discordant dunites are prominently cutting the concordant

358 lherzolites and dunites, especially around the central part of the ophiolite (Kadoshima &
359 Arai 2001). The younger discordant dunites are black in hand specimen, and seem
360 compact and hard on outcrop (Fig. 5c). This is in contrast to the concordant dunites that
361 are severely altered/weathered to be pale green in color and have been more strongly
362 eroded than the discordant ones due to mechanical weakness. This observation clearly
363 indicates different chemical and/or textural characteristics between the two types of
364 dunite. The discordant dunites are exclusively composed of olivine, highly serpentinized,
365 and euhedral to subhedral chromian spinel. The chromian spinel is opaque in thin
366 section and contains minute exsolution blebs of a Ti-rich phase (possibly Ti-rich
367 magnetite) (Fig. 5d). The olivine shows slightly lower Fo contents, 83 to 88, than in the
368 wall peridotite (89-90). The chromian spinel of this younger dunite is high in Cr# and
369 TiO₂, being around 0.6 and up to > 4 wt%, respectively (Fig. 6). It is low in both Cr#
370 and TiO₂ near the boundary with lherzolite (Fig. 6), suggesting fractional crystallization
371 (precipitation of minerals from the wall inward) or a reaction between the involved melt
372 and the lherzolite. The primary chromian spinel in the discordant dunite should have
373 contained higher TiO₂ contents before unmixing of the Ti-rich phase, being within the
374 chemical range of dunite spinels from oceanic hotspots (see Figs. 1 and 2). The magma

375 that produced the discordant dunite within the Lizard peridotite was of hotspot
376 (intra-plate) origin (Figs. 1, 2, 4 and 6). It was most probably tholeiitic (cf. Arai 1992).
377 The Lizard peridotite was, therefore, derived from the uppermost mantle that was
378 initially generated at a slow spreading ridge and was later impacted by an intra-plate
379 tholeiite magma.

380

381 CONCLUSIONS

382 The chromian spinel chemistry is highly useful to petrologically characterize ultramafic
383 plutonic rocks, especially dunitic rocks and chromitites, where chromian spinel is often
384 the only discriminating mineral. The trivalent cation ratio and TiO_2 content in chromian
385 spinel are important parameters in discrimination of the plutonic rocks in terms of
386 tectonic setting of formation. Discrimination diagrams based on spinel chemistry made
387 from plutonic rocks derived from well-constrained settings should be applied to
388 characterization of rocks from unknown origins. The spinel-based diagrams made for
389 volcanic rocks or magmas should not been used for discrimination of plutonic rocks in
390 tectonic setting of derivation, because chromian spinel shows different chemical ranges
391 between effusive and plutonic rocks even of the same magmatic affinity as discussed

392 above. Discrimination in Mg# of chromian spinel is sometimes unreliable, because the
393 Mg# in chromian spinel is strongly changeable depending on the thermal history in
394 olivine-rich rocks. For example, the chromian spinel in possible abyssal peridotites
395 suffered from low-temperature metamorphism at a subduction zone have lower Mg#s at
396 a given Cr# than abyssal peridotites from the present-day ocean floor, which are mostly
397 derived from near the spreading center and have not been cooled down sufficiently (e.g.
398 Hirauchi *et al.* 2008).

399

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406

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664 Figure Captions

665

666 Fig. 1. Relationships between TiO_2 contents and $\text{Cr}/(\text{Cr} + \text{Al})$ atomic ratios of
667 chromian spinels in plutonic rocks. Upper and lower panels are for dunites and other
668 possible cumulates, and for peridotites, respectively. (a) Mid-ocean ridges. Data
669 source: Allan and Dick (1996), Arai and Matsukage (1996), Cannat *et al.* (1997),
670 Dick (1989), Dick and Natland (1996), Fujii (1990), Niida (1997), Komor *et al.*
671 (1990), Prinz *et al.* (1976), and Tartarotti *et al.* (2002). (b) Arcs. Data source: Abe *et*
672 *al.* (1992, 1995), Arai *et al.* (2004), Barsdell and Smith (1989), Bloomer and
673 Hawkins (1983), Bloomer and Fisher (1987), Conrad and Kay (1984), Debari *et al.*
674 (1987), Delong *et al.* (1975), Ishii (1985), Ishii *et al.* (1992, 2000), Ishimaru (2004),
675 Ninomiya and Arai (1992), Ohara and Ishii (1998), and Yamamoto (1984). (c)
676 Oceanic hotspots. Data source: Clague (1988), Sen and Presnall (1986), Tracy
677 (1980), and Tanaka (1999). Peridotites, lherzolites and harzburgites. Pl, plagioclase.
678 Ol, olivine. Note the different compositional ranges between the three tectonic
679 settings. Scales of vertical axis are different between (a) and (b), and (c).

680

681 Fig. 2. Cr-Al-Fe³⁺ atomic relationships of chromian spinels in plutonic rocks. Right
682 and left panels are for dunites and other possible cumulates, and for peridotites,
683 respectively. (a) Mid-ocean ridges. (b) Arcs. (c) Oceanic hotspots. Peridotites,
684 lherzolites and harzburgites. Pl, plagioclase. Ol, olivine. Data source as in Fig. 1.
685

686 Fig. 3. Relationships between Mg/(Mg + Fe²⁺) and Cr/(Cr + Al) atomic ratios of
687 chromian spinels in plutonic rocks. Upper and lower panels are for dunites and other
688 possible cumulates, and for peridotites, respectively. (a) Mid-ocean ridges. (b) Arcs.
689 (c) Oceanic hotspots. Peridotites, lherzolites and harzburgites. Pl, plagioclase. Ol,
690 olivine. Data source as in Fig. 1.
691

692 Fig. 4. Relationships between TiO₂ contents and Fe³⁺/(Cr + Al + Fe³⁺) atomic ratios
693 of chromian spinels in plutonic rocks. Upper and lower panels are for dunites and
694 other possible cumulates, and for peridotites, respectively. (a) Mid-ocean ridges. (b)
695 Arcs. (c) Oceanic hotspots. Peridotites, lherzolites and harzburgites. Pl, plagioclase.
696 Ol, olivine. The fields for MOR (a) and hotspot (c) plutonics are shown in the panel
697 (b). The field for MORB spinels (Arai, 1992) is shown in the panel (a) for

698 comparison. Data source as in Fig. 1. Fields for the main clusters of mid-ocean ridge
699 (MOR) and hotspot spinels are shown in the panel (b).

700

701 Fig. 5. Photographs of discordant dunites. (a) Outcrop of discordant dunites (D)
702 within foliated mantle harzburgite from Wadi Rajmi, the northern Oman ophiolite.
703 (b) Photomicrograph of a partially serpentinized discordant dunite from Wadi Rajmi.
704 Plane-polarized light. (c) Outcrop of a discordant dunite (D; selectively eroded)
705 from the Lizard ophiolite. (d) Photomicrograph of a chromian spinel grain with
706 high-Ti exsolution blebs (brighter) in a partially serpentinized discordant dunite
707 from the Lizard ophiolite. Reflected plane-polarized light. Note the bright band
708 fringing the right-side margin is remnants of carbon coating.

709

710 Fig. 6. Chromian spinel compositions in discordant dunites from the Oman and
711 Lizard ophiolites. Note the different compositional characteristics between the two
712 dunite spinels. (a) TiO_2 vs. $\text{Cr}/(\text{Cr} + \text{Al})$ atomic ratio. Compare with Figure 1. (b)
713 TiO_2 vs. $\text{Fe}^{3+}/(\text{Cr} + \text{Al} + \text{Fe}^{3+})$ atomic ratios. Compare with Figure 4. (c) Cr-Al- Fe^{3+}
714 atomic ratios. The fields for MOR and hotspot dunites (Figure 2) are shown in the

715 panel (b).

716

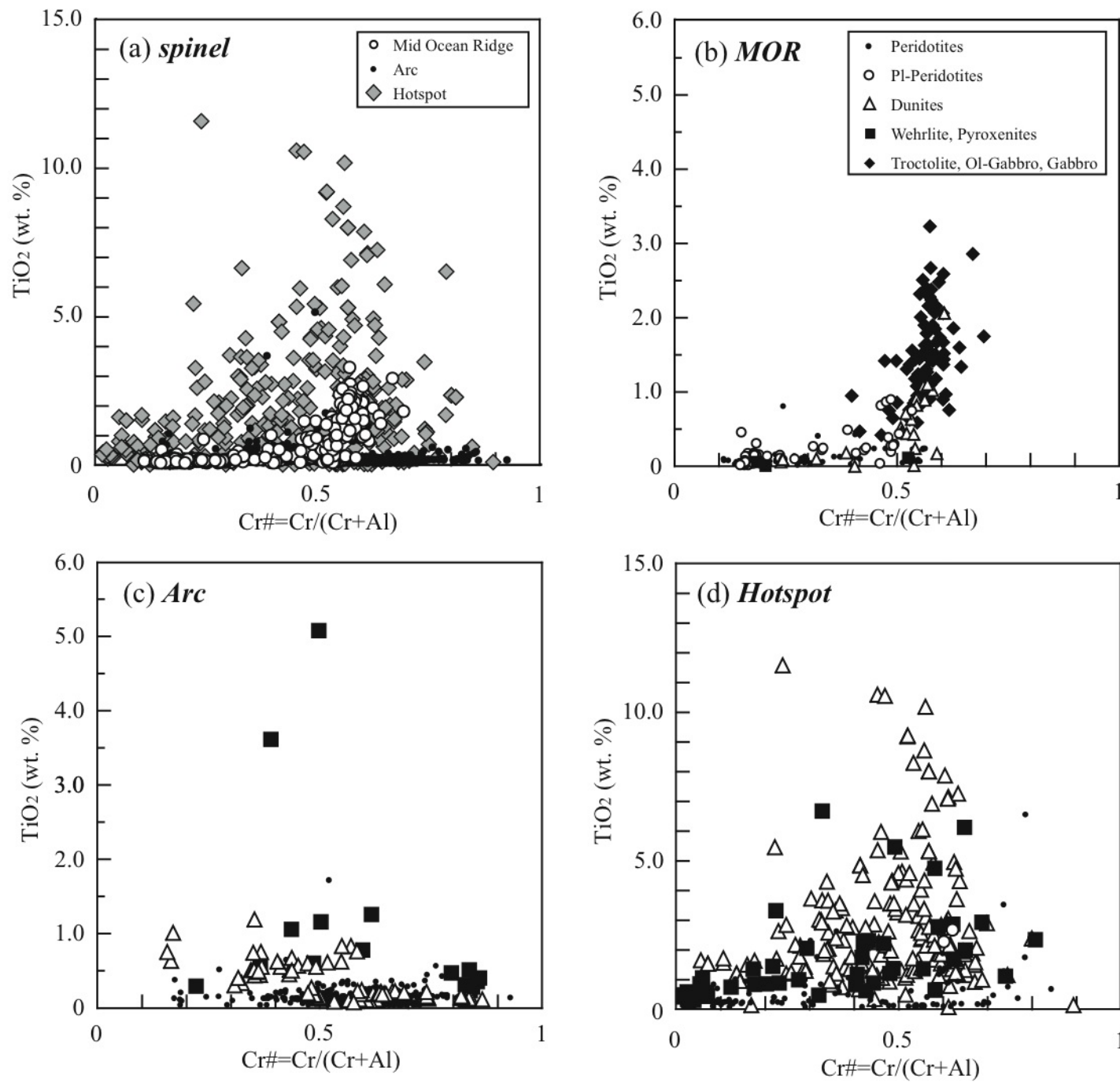


Fig.1 Arai et al.

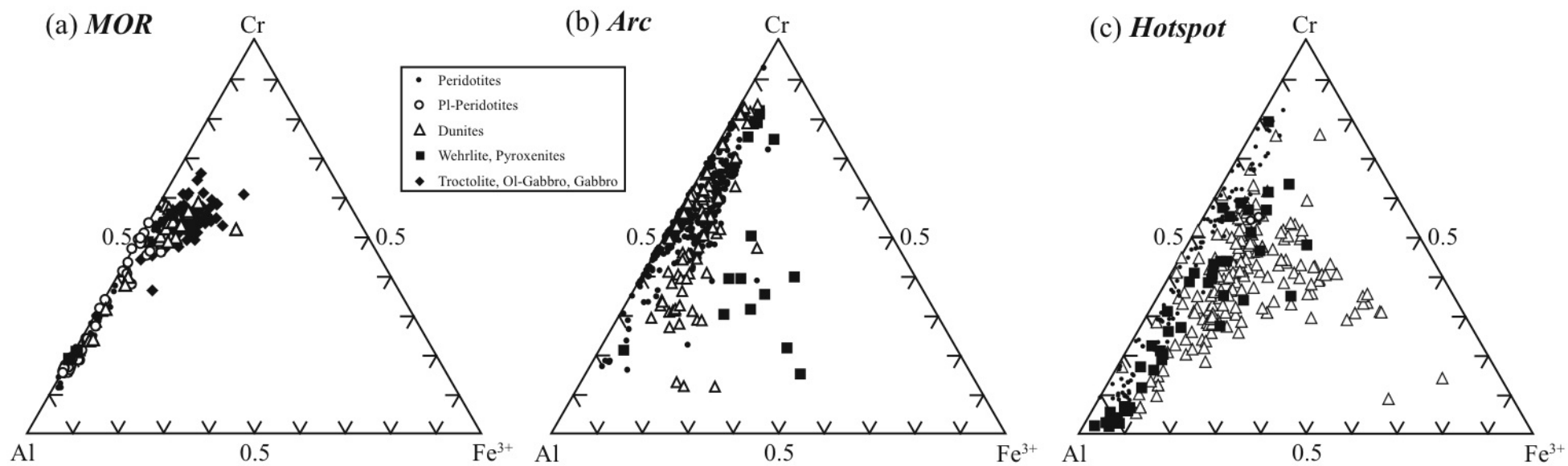


Fig. 2 Arai et al.

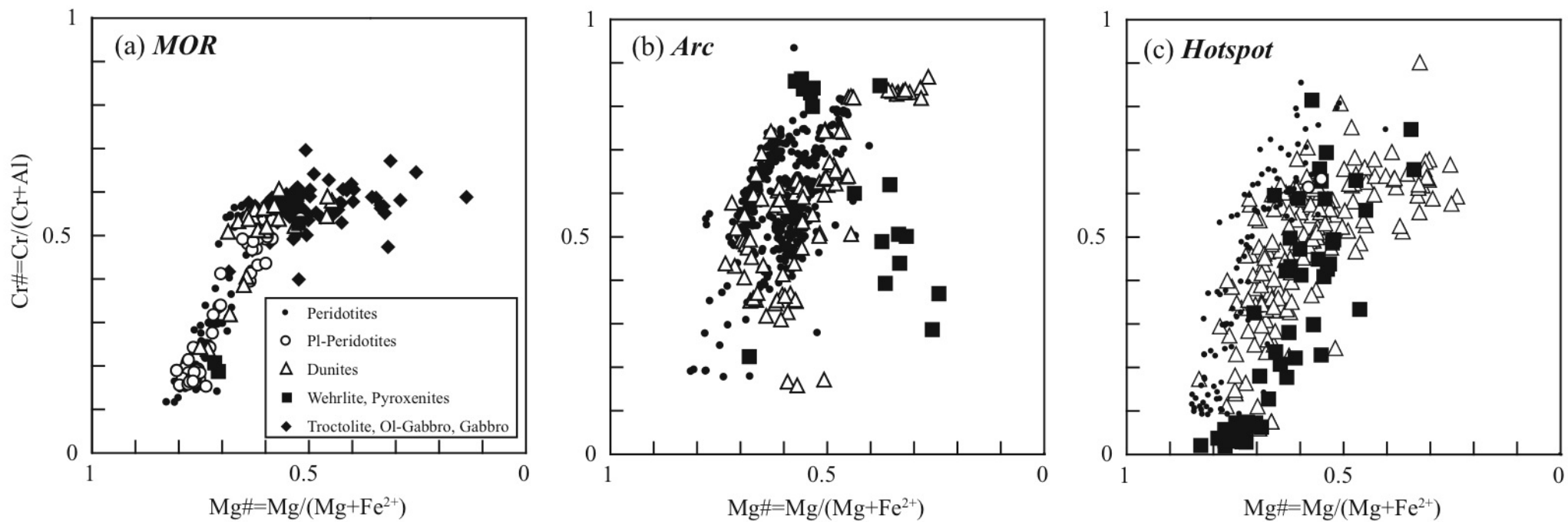


Fig. 3 Arai et al.

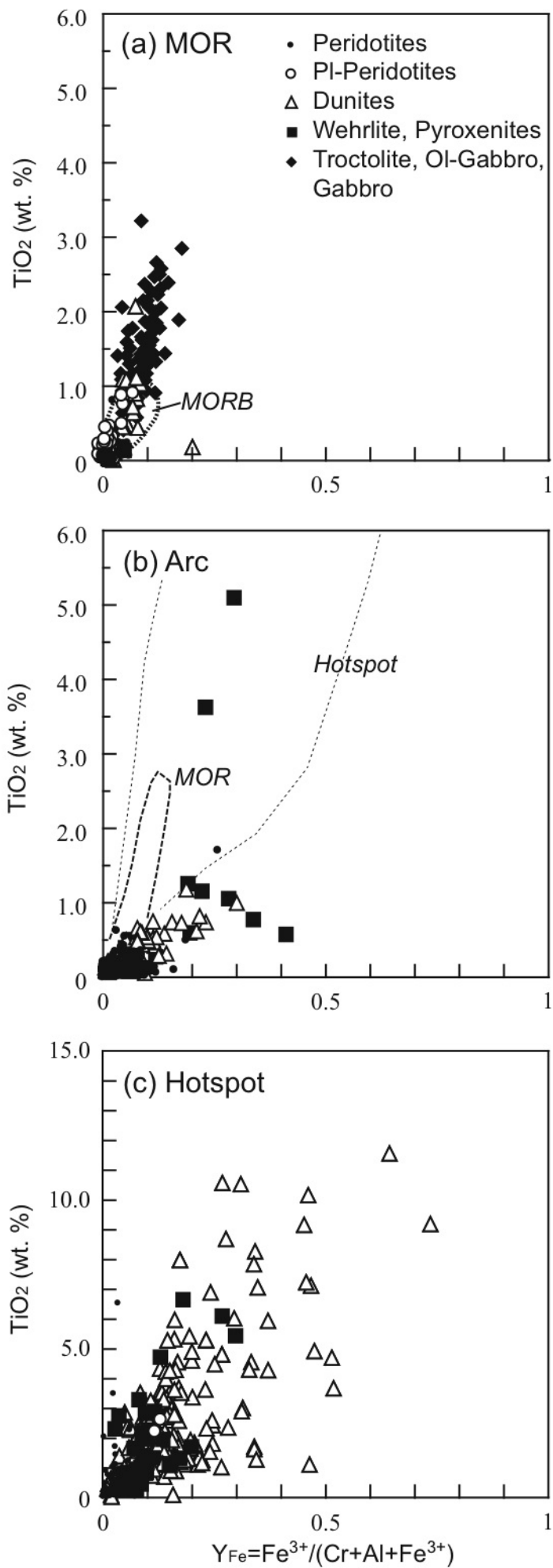


Fig. 4 Arai et al.

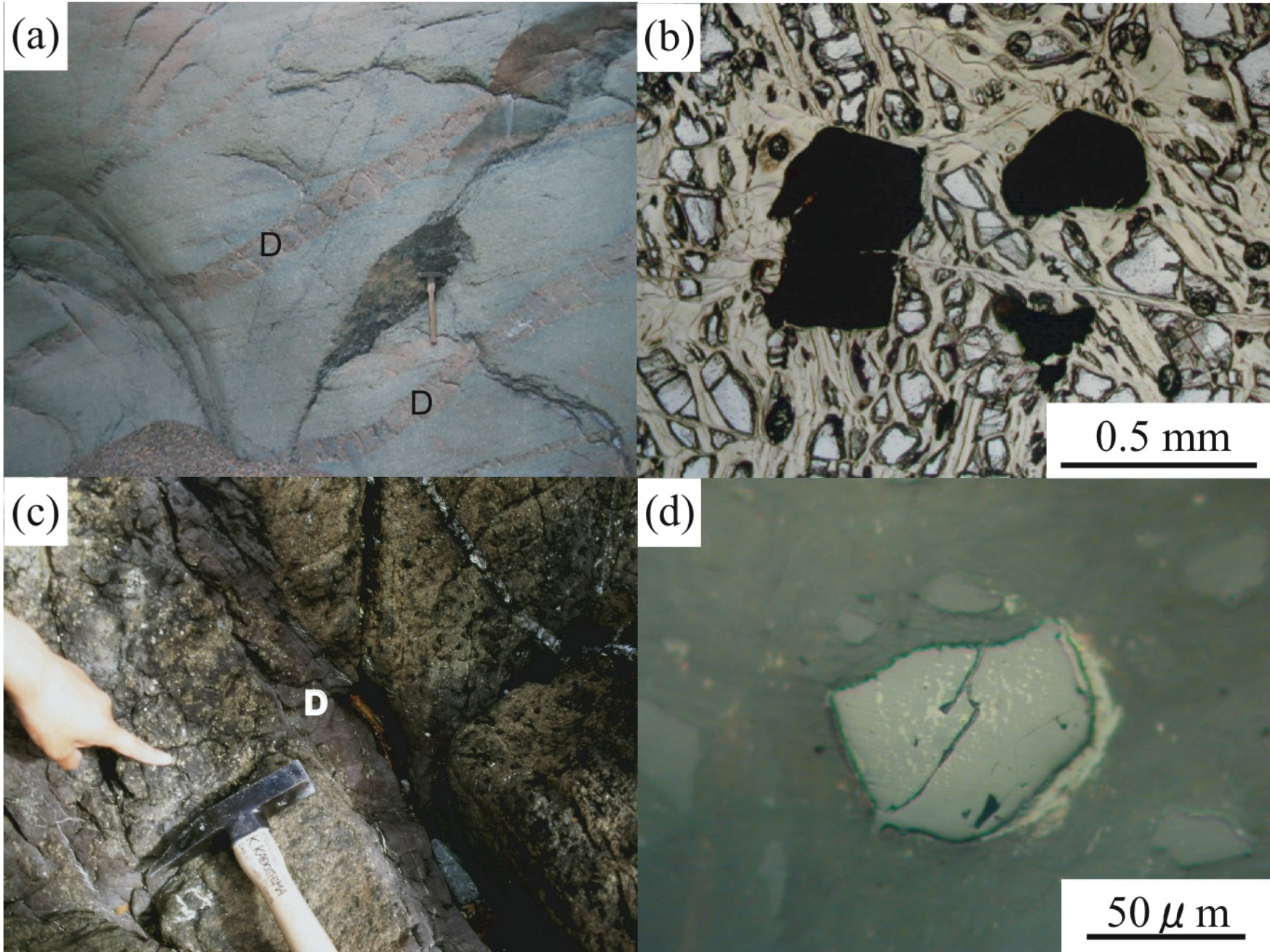


Fig. 5 Arai et al.

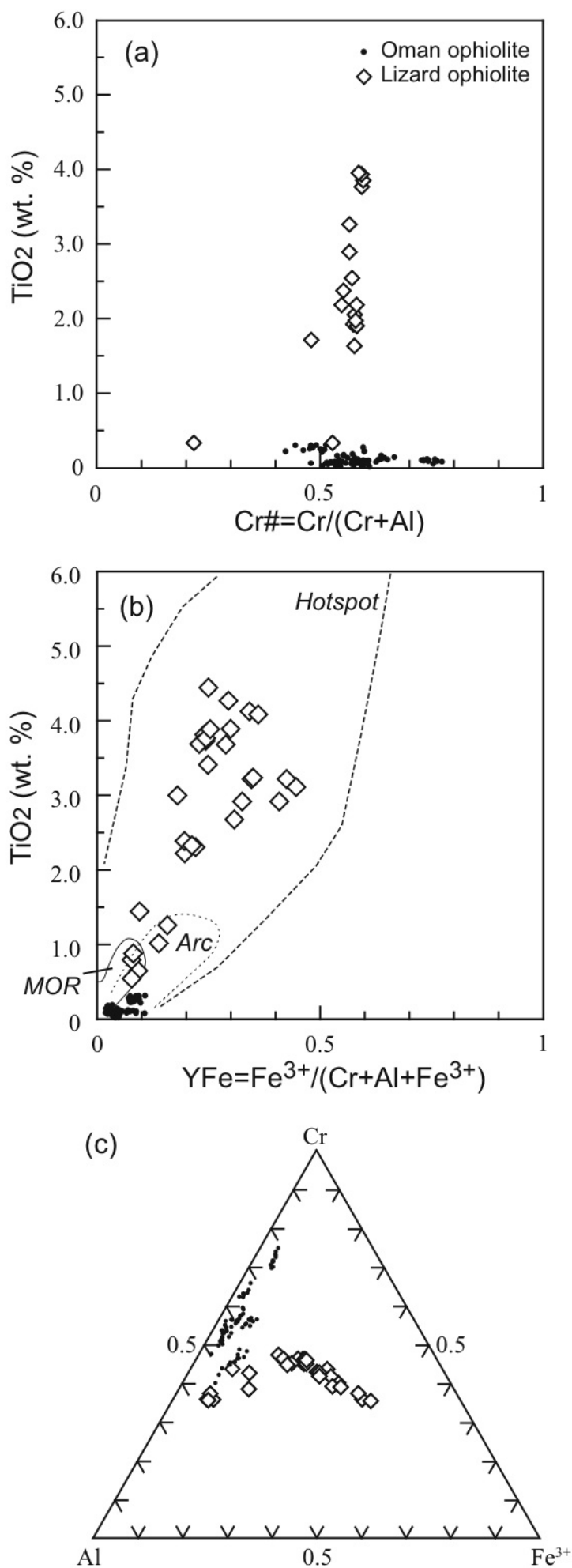


Fig. 6 Arai et al.